

Educating Kids & Exciting Teachers about Science:
A Model from the Plasma Science Community

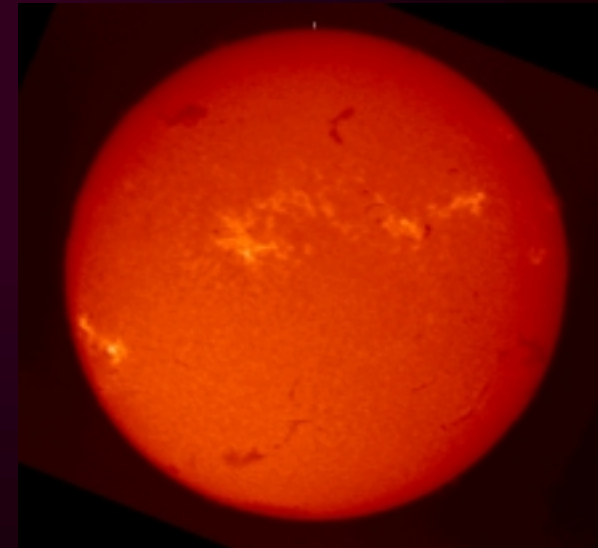
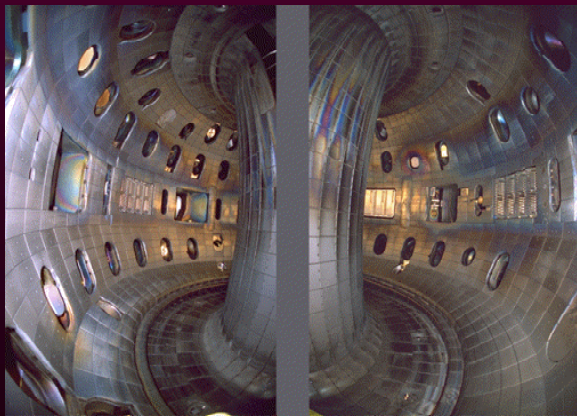


The Power of the Universe on Earth: Plasma Physics and Fusion Energy

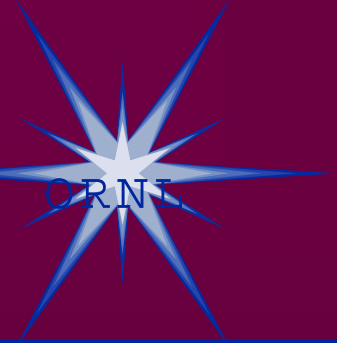
David Newman

Physics Department
University of Alaska - Fairbanks

What is a plasma?
Why should we care?
Where are the questions?



March 14, 2000



Science and Education



“...if you tell me I listen ... if you teach me I learn ... if you involve me I remember ...”

(Jim Diaz quoting Ben Franklin)

Outreach and education are intrinsically linked

Stimulate interest and you will stimulate learning

Scientists at all levels must get involved in sharing their science



Plasma Science in Education and Outreach



Elementary schools

High schools

University/College

Teacher training

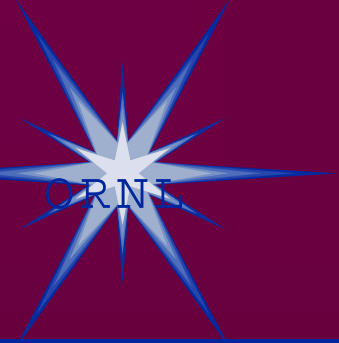
General Outreach

In all of these settings:

Outreach, materials development and
direct educational efforts are underway



1999 APS/DPP Outreach/Expo
Courtesy of Carol Danielson



Outline



What is a plasma?

Where do we find them?

Why are we interested in them?

Astrophysics

Plasmas all around us

Fusion energy

More on fusion energy.

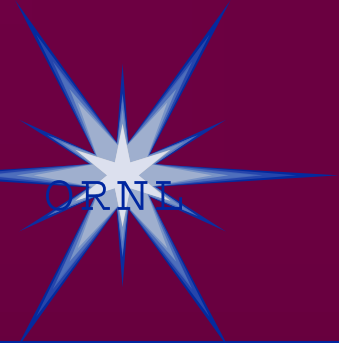
Charged particles moving in a magnetic field.

Turbulence



Picture courtesy of Jan Curtis

<http://climate.gi.alaska.edu/Curtis/curtis.htm>



What is a plasma?



A plasma is an ionized gas.

Plasma is called the “**fourth state of matter.**”

More than 99% of the known mass of the universe is in the plasma state.

‘Plasma’ was coined by Tonks and Langmuir in (1929):

“...when the electrons oscillate, the positive ions behave like a rigid jelly...”



Where do we find plasmas?



Examples of plasmas on Earth:

Lightning

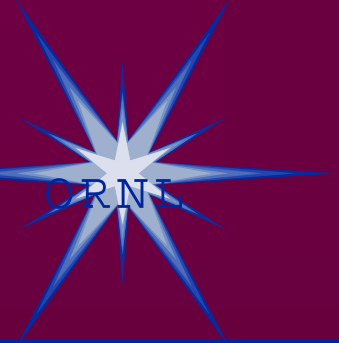
Neon and Fluorescent Lights

Laboratory Experiments

Examples of astrophysical plasmas:

The sun and the solar wind

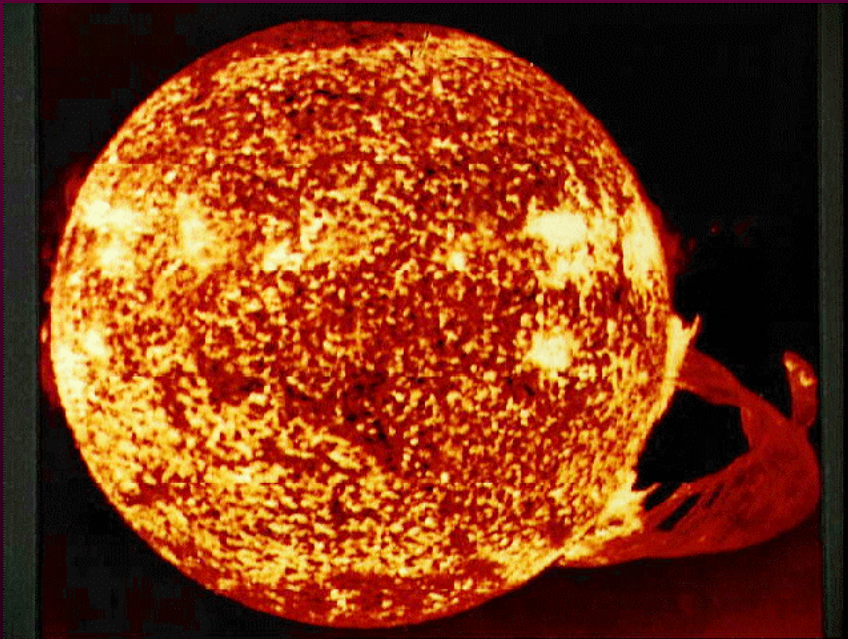
Stars, interstellar medium



Astrophysical plasmas

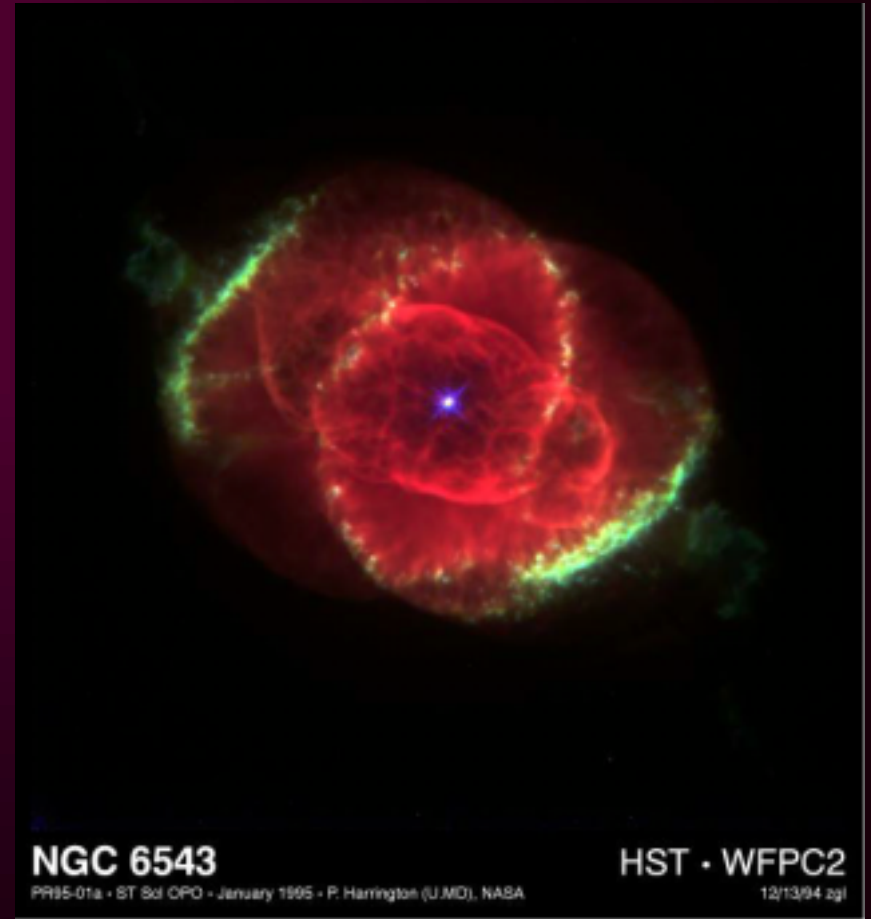


The Sun

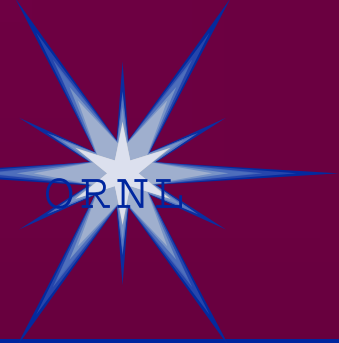


<http://bang.lanl.gov/solarsys/>

Catseye
Nebula



<http://www.stsci.edu:80/>



Astrophysical Plasmas

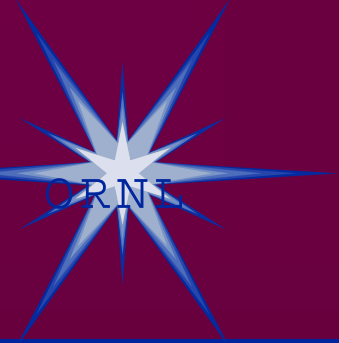


MyCn18, a young planetary nebula located about 8,000 light-years away



Pictures courtesy of NASA
http://nssdc.gsfc.nasa.gov/photo_gallery/

Galaxy NGC 4414, is 19.1 megaparsecs or about 60 million light-years



Plasmas on Earth



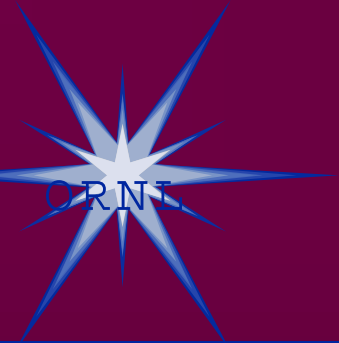
Laboratory Experiments



<http://FusEdWeb.pppl.gov/>

Lightning





Why are we interested in plasmas?



Fusion Energy

Potential source of safe, abundant energy.

Astrophysics

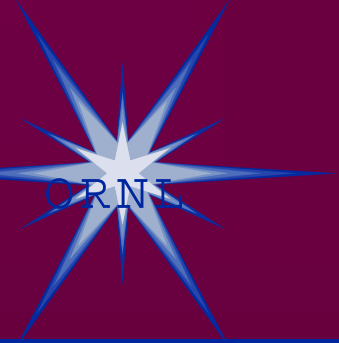
Understanding plasmas helps us understand stars, stellar evolution and the evolution of the Universe.

Upper atmospheric dynamics

The upper atmosphere is an important plasma.

Plasma Applications

Plasmas can be used to build computer chips and to clean up toxic waste.



Plasmas around us

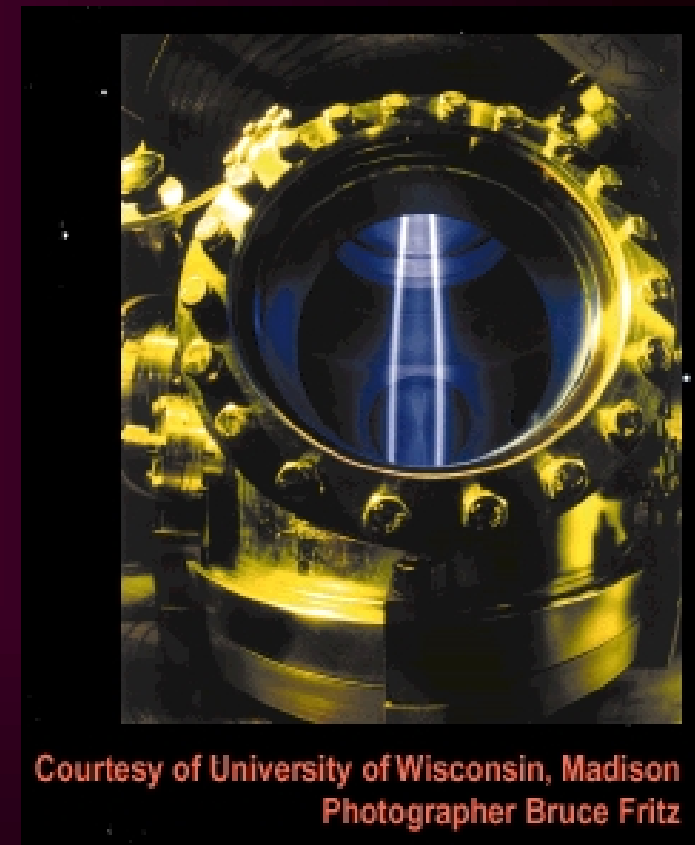
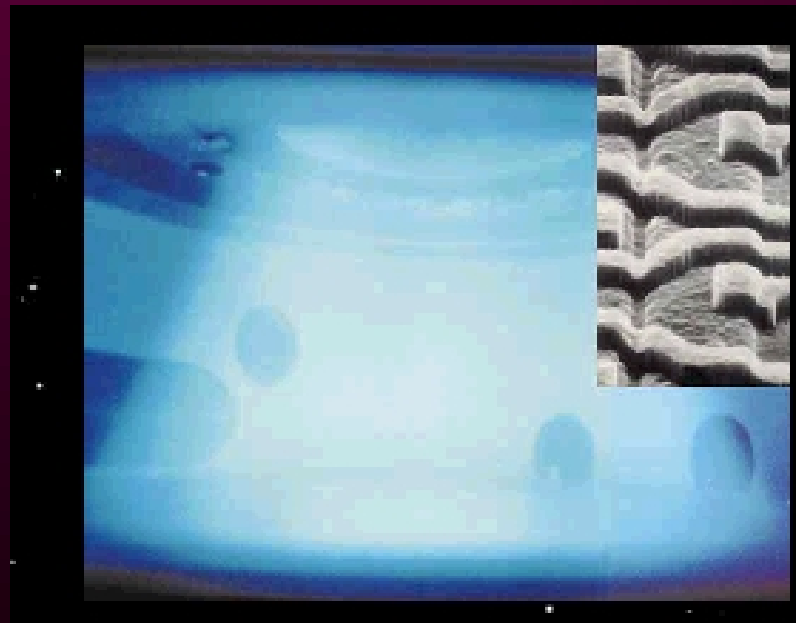
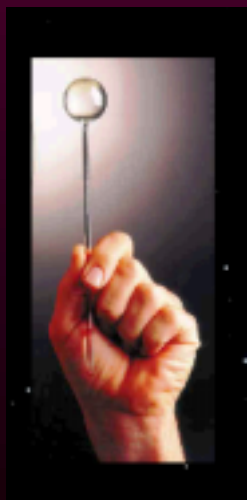


Plasma processing

Computer chips

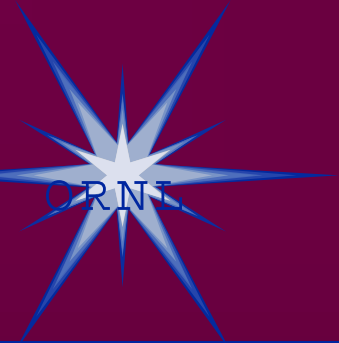
Surface modification

Lighting



Courtesy of University of Wisconsin, Madison
Photographer Bruce Fritz

Pictures from
Plasmas: The 4th State of Matter



Plasmas around us



Plasma displays
Decontamination
Plasma Thrusters



Plasma torch from the
Plasma Technology Research Center



Electrograph Plasma Display
<http://www.electrograph.com/detail.asp>

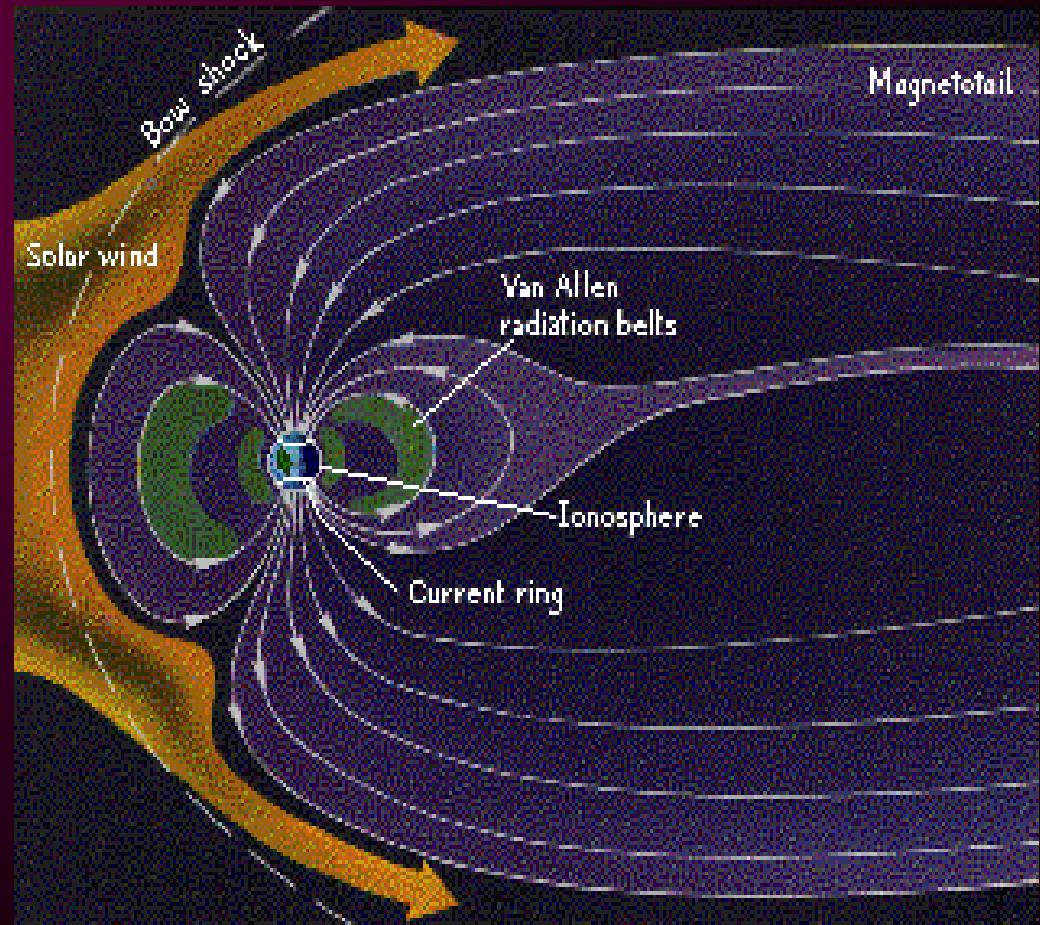
ORNL

The solar wind (a plasma) interacts with the Earth's magnetic field



The sun emits mass in the form of plasma at velocities of up to 500 km/s.

This solar wind causes the Earth's magnetic field to compress creating a shock wave called the Bow wave.



From Stars, James Kaler



Interactions between the earth's magnetic field and a plasma can have spectacular results



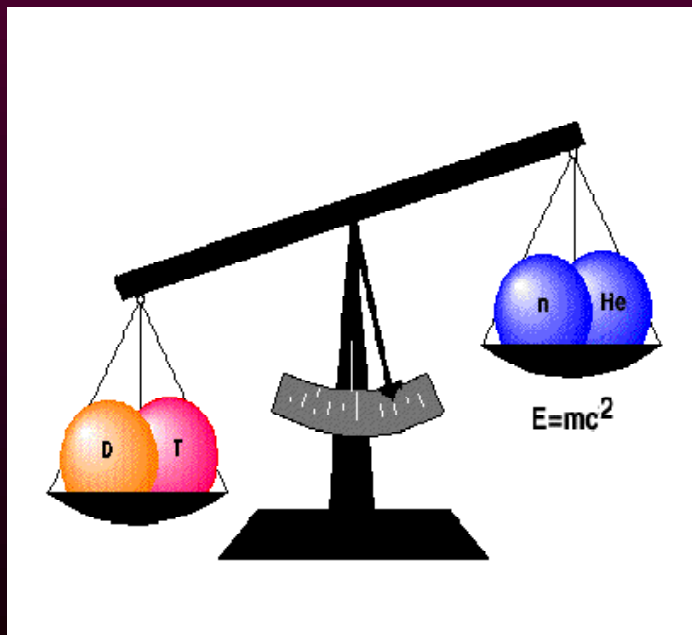
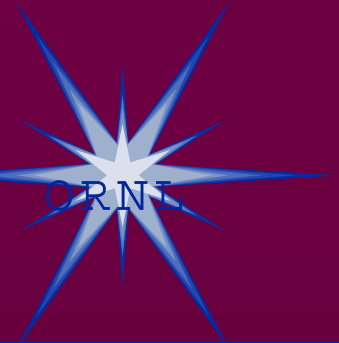
The northern lights
(aurora borealis)



Photo by David Fritts

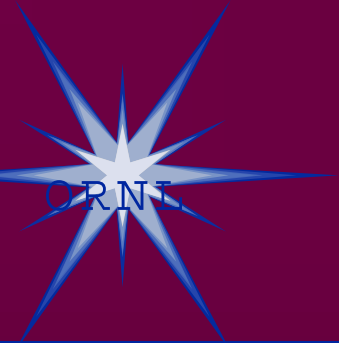
<http://dac3.pfrr.alaska.edu:80/~pfrr/AURORA/INDEX.HTM>

Mass goes into energy in a fusion reaction



Reactants	Fusion	Products
D	20 keV	3.5 MeV 4He
T	20 keV	14.1 MeV n

A diagram illustrating a fusion reaction. On the left, two reactants are shown: a deuterium nucleus (D) with one white and one yellow sphere, and a tritium nucleus (T) with one white and two yellow spheres. In the center, a large red starburst represents the fusion process. On the right, the products are shown: a helium-4 nucleus (4He) with two white and two yellow spheres, and a neutron (n) with one white sphere. Orange arrows point from the reactants towards the fusion point and from the fusion point towards the products.



Properties of plasmas

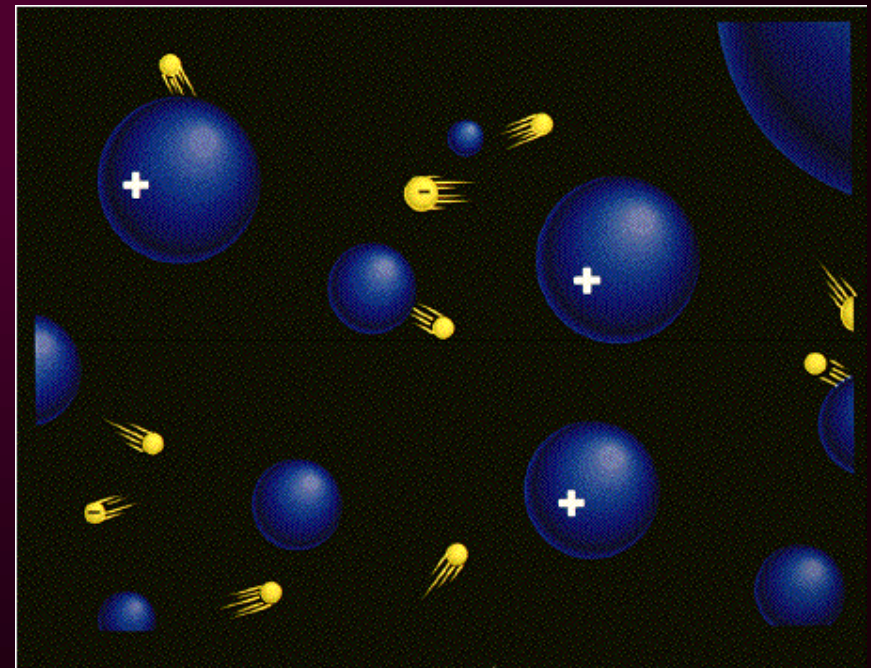


A collection of positively and negatively charged particles.

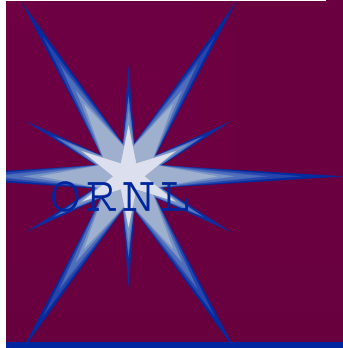
Plasmas interact strongly with electric and magnetic fields.

Plasmas support many different types of waves and oscillations.

Cartoon of a plasma



<http://demo-www.gat.com/>



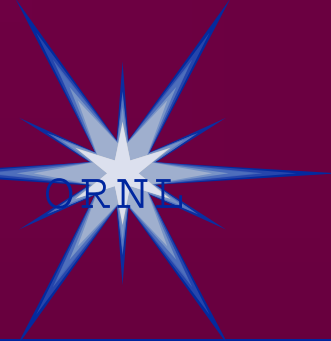
More on Fusion Energy



Much of plasma physics research has been motivated by the goal of controlled fusion energy.

Fusion energy is a form of nuclear energy which is emitted when two light nuclei combine to form a single more stable nuclei.

The sun and stars derive their energy from fusion.



Why is Fusion power needed?



Country	Consumption (kW-h/capita)
US	12000
Developed World Avg.	6000
World Avg.	1500
China	500
India	250

1990 Energy use per capita

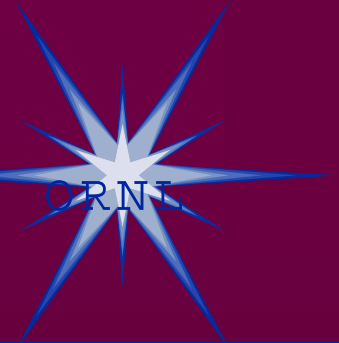
•Projected change in consumption by increasing to world average

Country	Energy Use 1990 (GW)	Energy Use 2020 (GW)
China	120	500
India	65	450

•If fossil Catastrophe Looms

For more information see:

http://www.foe.er.doe.gov/More_HTML/Artsimovich/PKKawPaper.html



Fuel and waste products

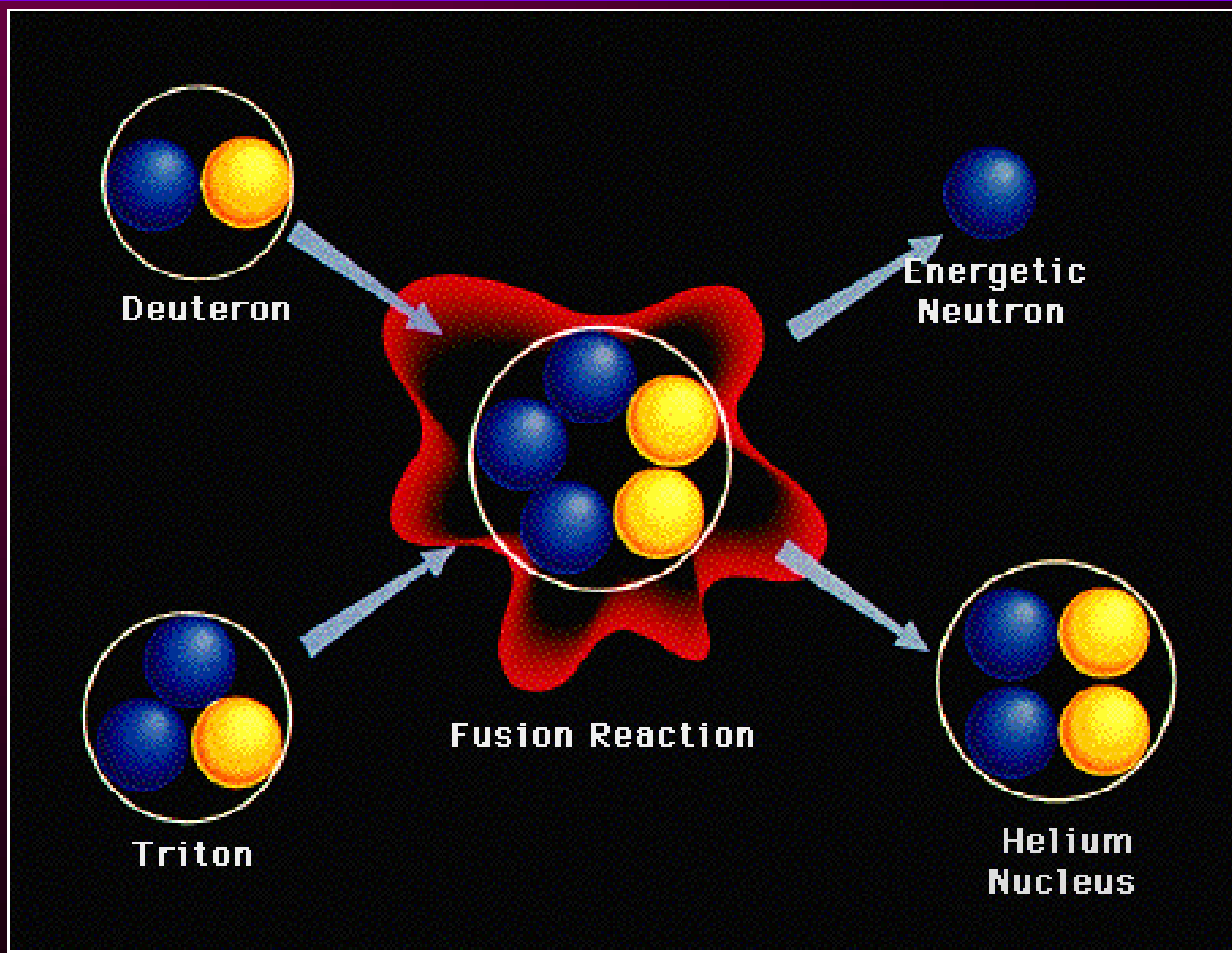
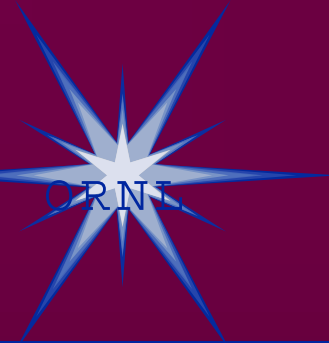


Fuel and waste for coal plants(most readily available energy source) vs D-T fusion plant

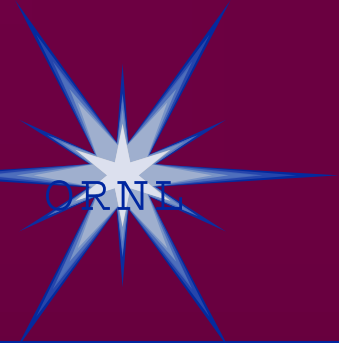
DAILY FUEL CONSUMPTION DAILY WASTE PRODUCTION 1,000 MEGAWATTS		
	COAL PLANT	D-T FUSION PLANT
F U E L	9,000 T. COAL	1.0 LB D ₂ 3.0 LB Li ⁶ (1.5 LB T ₂)
W A S T E	30,000 T. CO ₂ 600 T. SO ₂ 80 T. NO ₂	4.0 LB He ⁴

<http://www.pppl.gov>

Deuterium and tritium combine to form helium, a neutron and fusion energy.



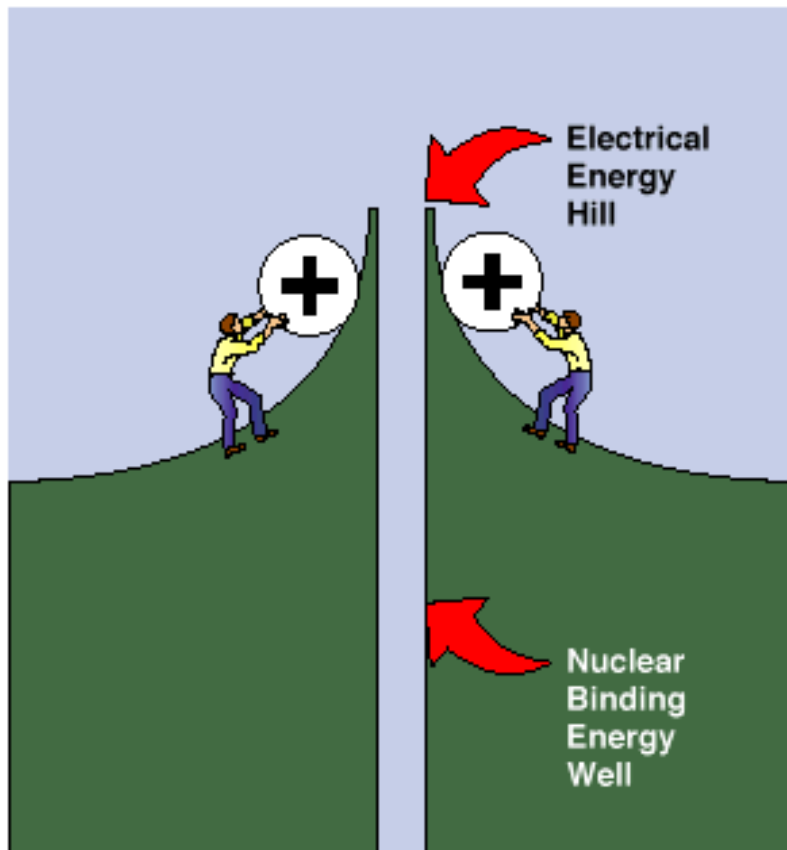
<http://FusEdWeb.pppl.gov/>



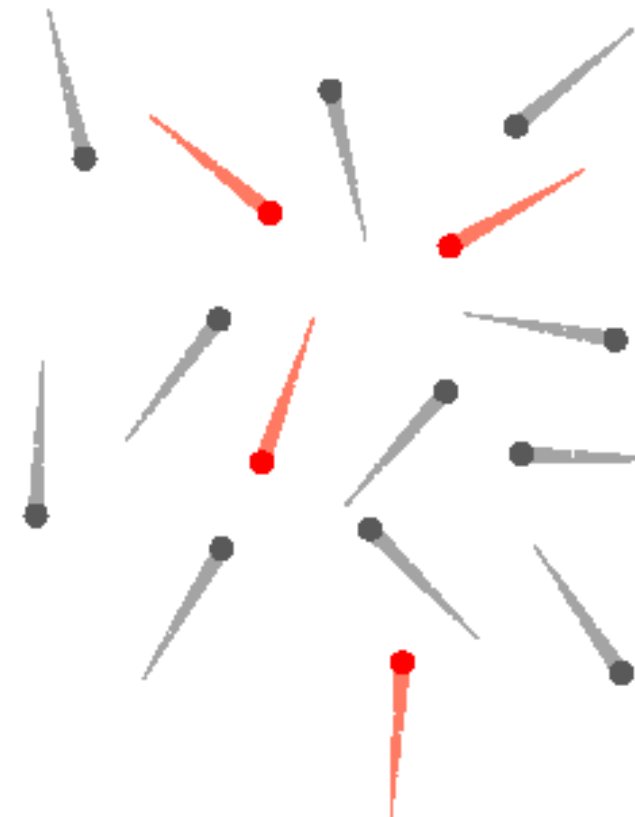
High temperatures and densities are needed



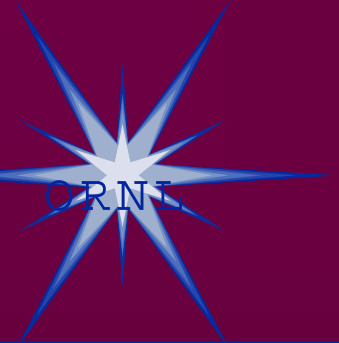
High temperature



Will they meet?



- ✓ Long confinement time
- ✓ Increased pressure

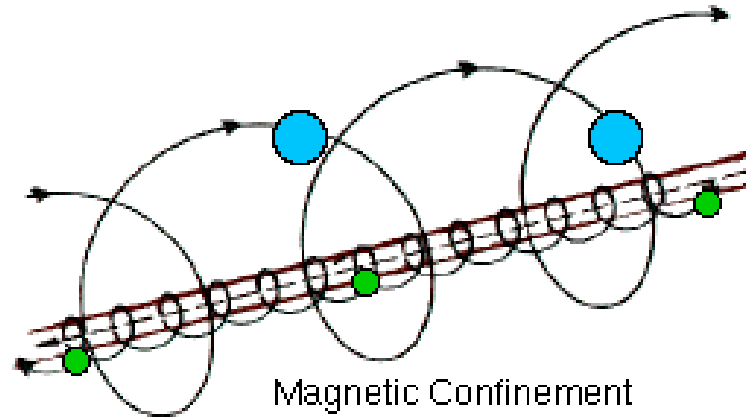


Methods for confinement

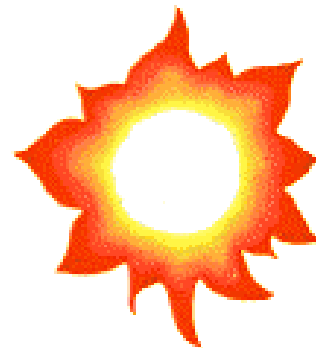


Hot plasmas are confined with gravitational fields in stars.

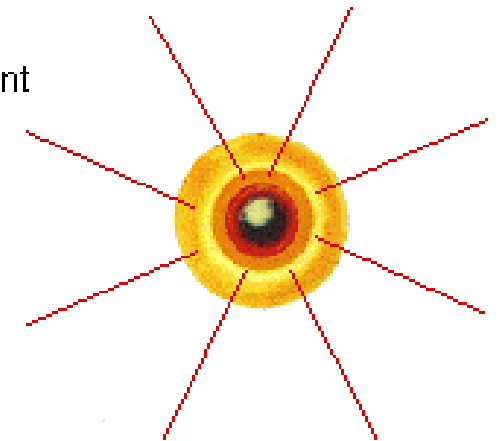
In fusion energy experiments magnetic fields and lasers are used to confine the hot plasma.



Magnetic Confinement




Gravitational Confinement in the Sun and Stars



Inertial Confinement Using Lasers

<http://FusEdWeb.pppl.gov/>



What must be achieved to obtain fusion energy?



Contain a high temperature, T , high density, n , plasma for a long enough time, τ , to achieve ignition (power out \gg power in).

A measure of plasma performance is thus given by:

$$nT\tau$$

density \times temperature \times confinement time

Two major approaches to fusion (D-T)



Magnetic confinement

Temperature $\approx 10^8$ °C (10 keV)

$\eta\tau \approx 10^{15}$ Atoms ·seconds / cm³

$\tau \approx 10$ seconds (magnetic “bottle”)

$\eta \approx 10^{14}$ Atoms / cm³ (10⁻⁵ times the density of air)

Inertial confinement

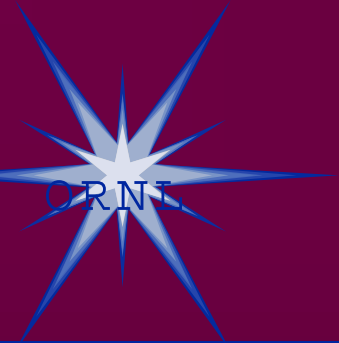
Temperature $\approx 10^8$ °C (10 keV)

$\eta\tau \approx 10^{15}$ Atoms ·seconds / cm³

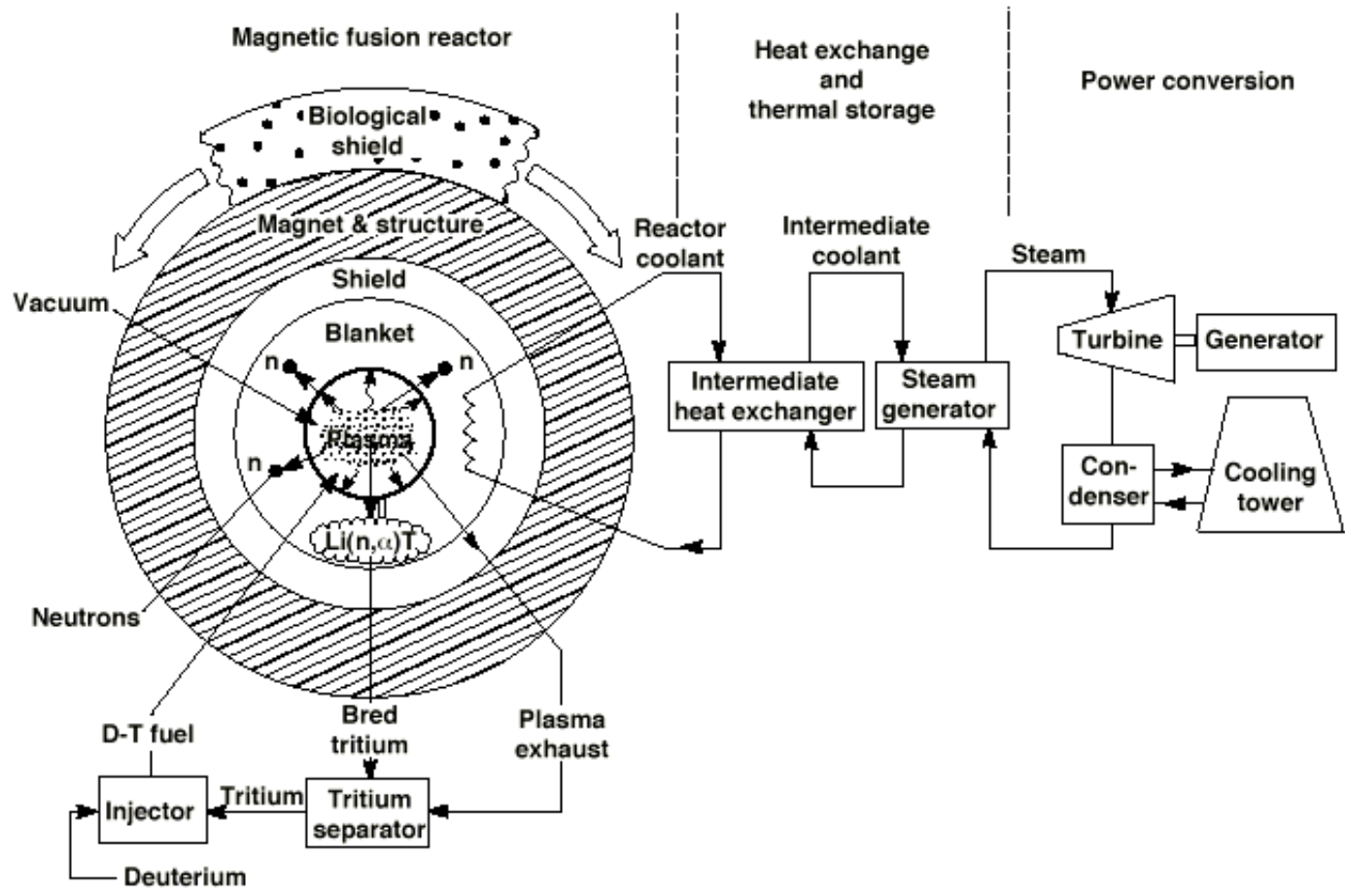
$\tau \approx 3 \times 10^{-11}$ seconds (microexplosion, inertial “bottle”)

$\eta \approx 3 \times 10^{25}$ Atoms / cm³ (12 times the density of lead!

~ 1000 times the density of liquid DT!)



Power Plant Schematic



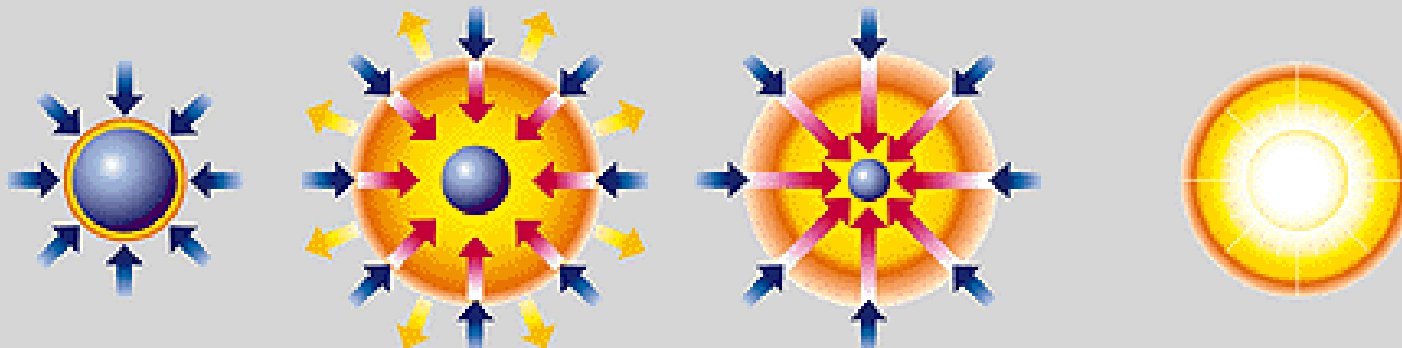
05-00-0398-0400
©2005 ORNL

The Inertial Confinement Fusion Concept

→ Laser energy

→ Blowoff

→ Inward transported thermal energy



**Atmosphere
formation**

Laser beams rapidly heat the surface of the fusion target forming a surrounding plasma envelope.

Compression

Fuel is compressed by the rocket-like blowoff of the hot surface material.

Ignition

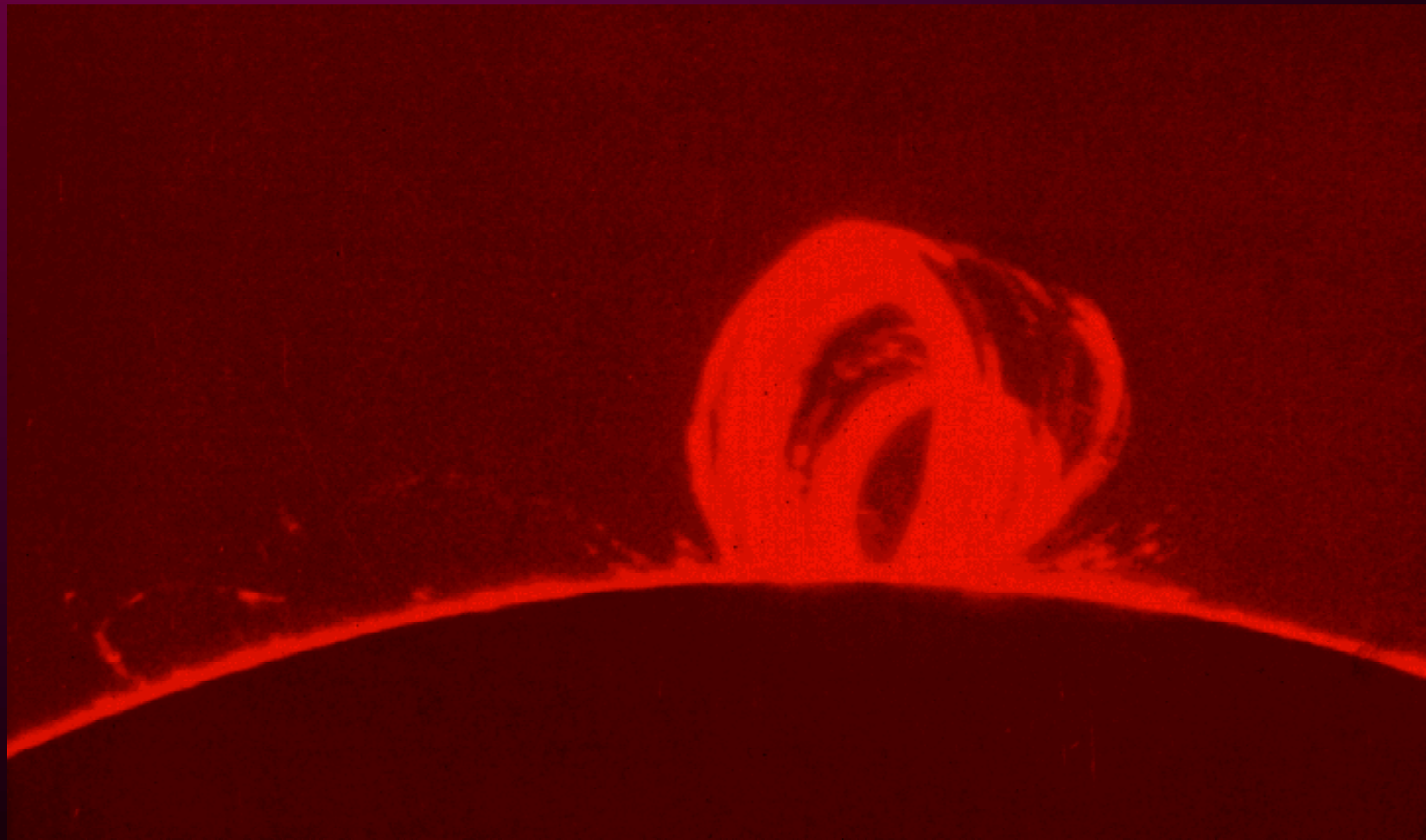
During the final part of the laser pulse, the fuel core reaches 20 times the density of lead and ignites at 100,000,000°C.

Burn

Thermonuclear burn spreads rapidly through the compressed fuel, yielding many times the input energy.

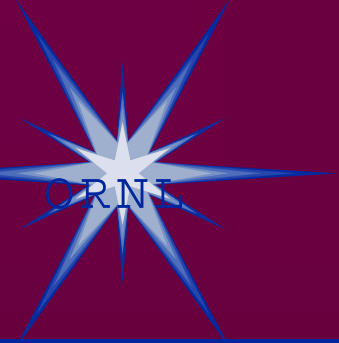


Particles in a Magnetic field



Picture courtesy of NASA

http://nssdc.gsfc.nasa.gov/photo_gallery/

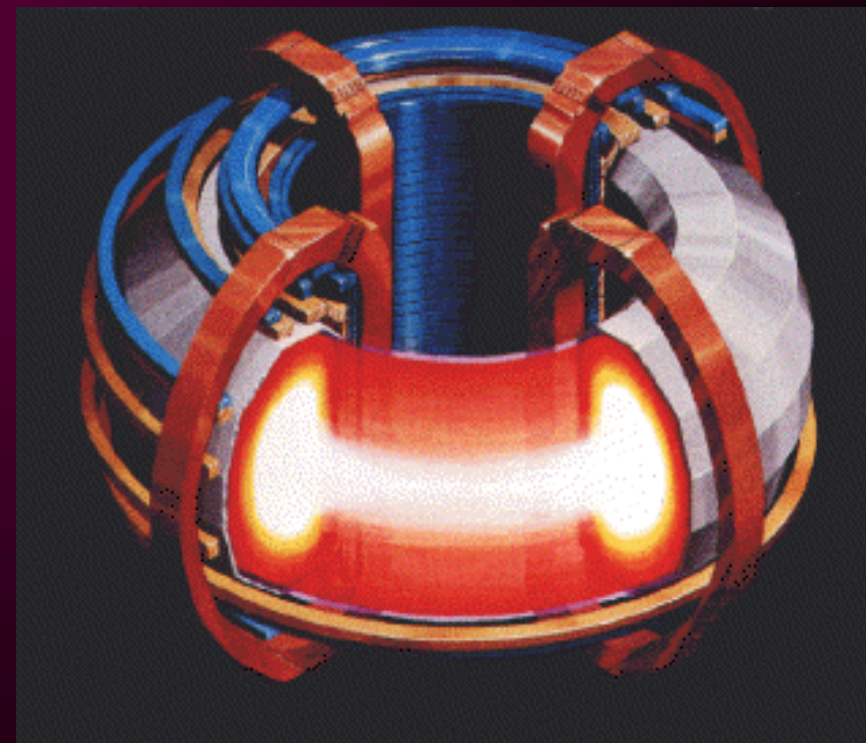


Controlling fusion with magnetic fields

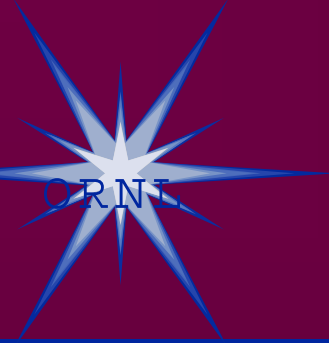


Most magnetic confinement devices in use today have a toroidal shape.

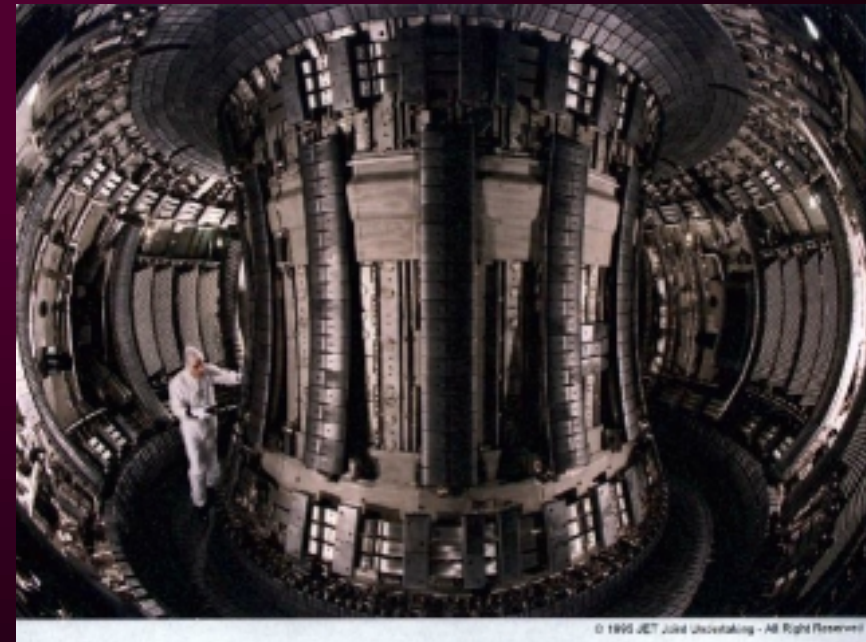
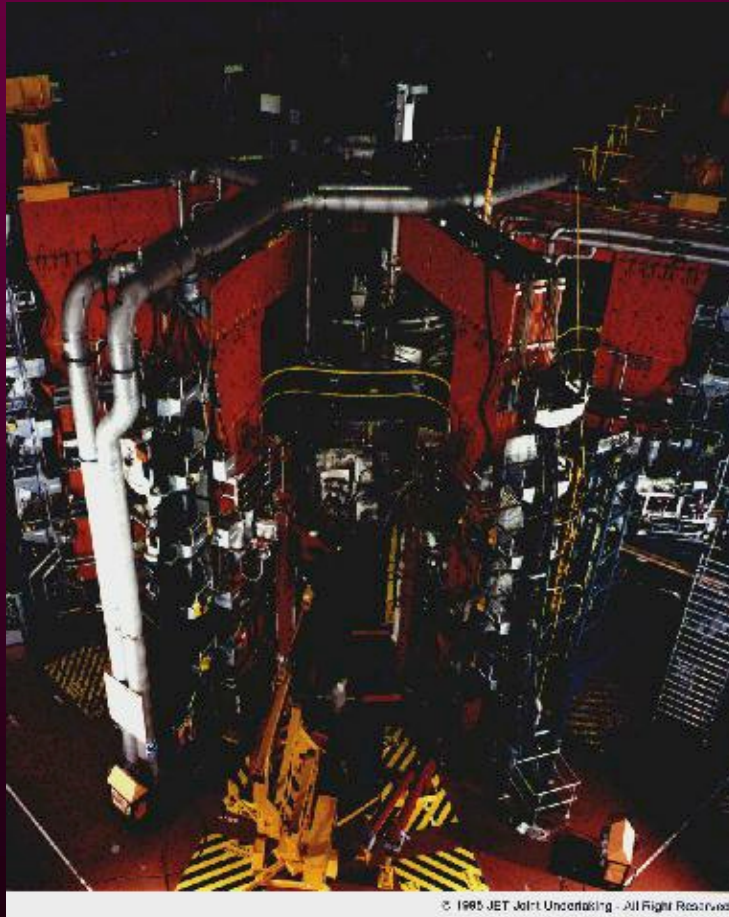
Large magnetic fields are created by driving currents through coils wrapped around the torus.



<http://demo-www.gat.com/>



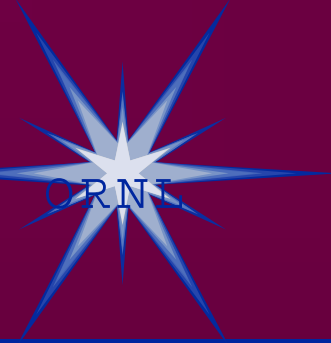
Joint European Torus: the largest confinement device ever built



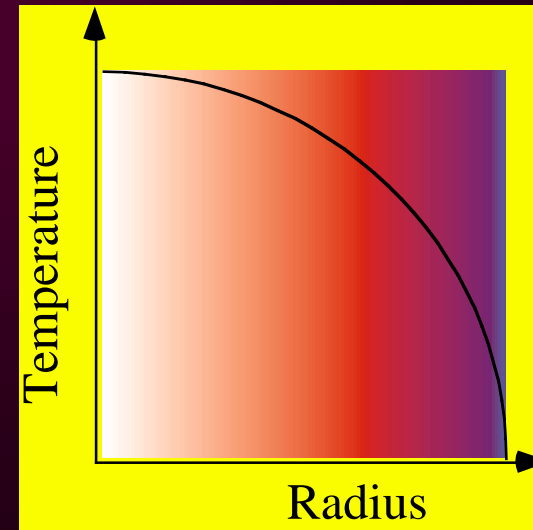
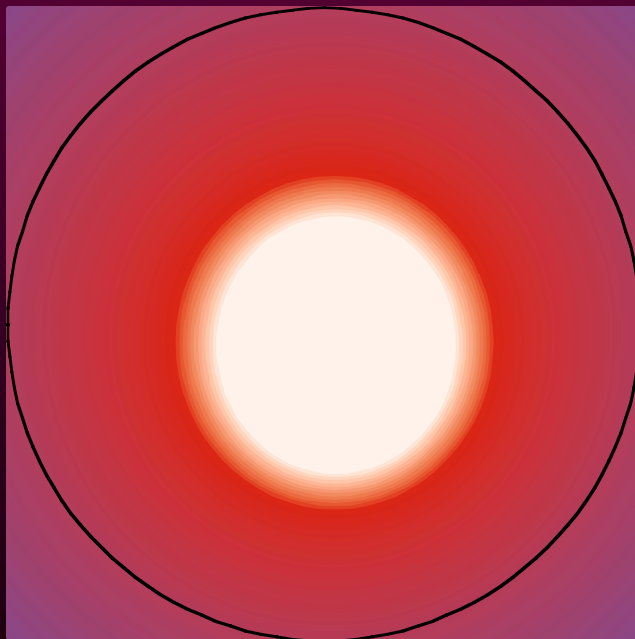
<http://www.jet.uk/>

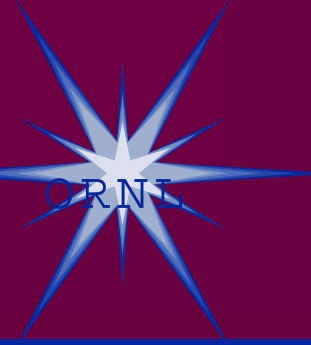
<http://www.jet.uk/>

Need to control temperature and density



We need the core hot enough for fusion, yet the edge cool enough not to melt the walls





But nature abhors gradients



Whenever a slope (gradient) gets too steep,
nature finds a way to flatten it out

Mountains get eroded

sand and snow avalanche

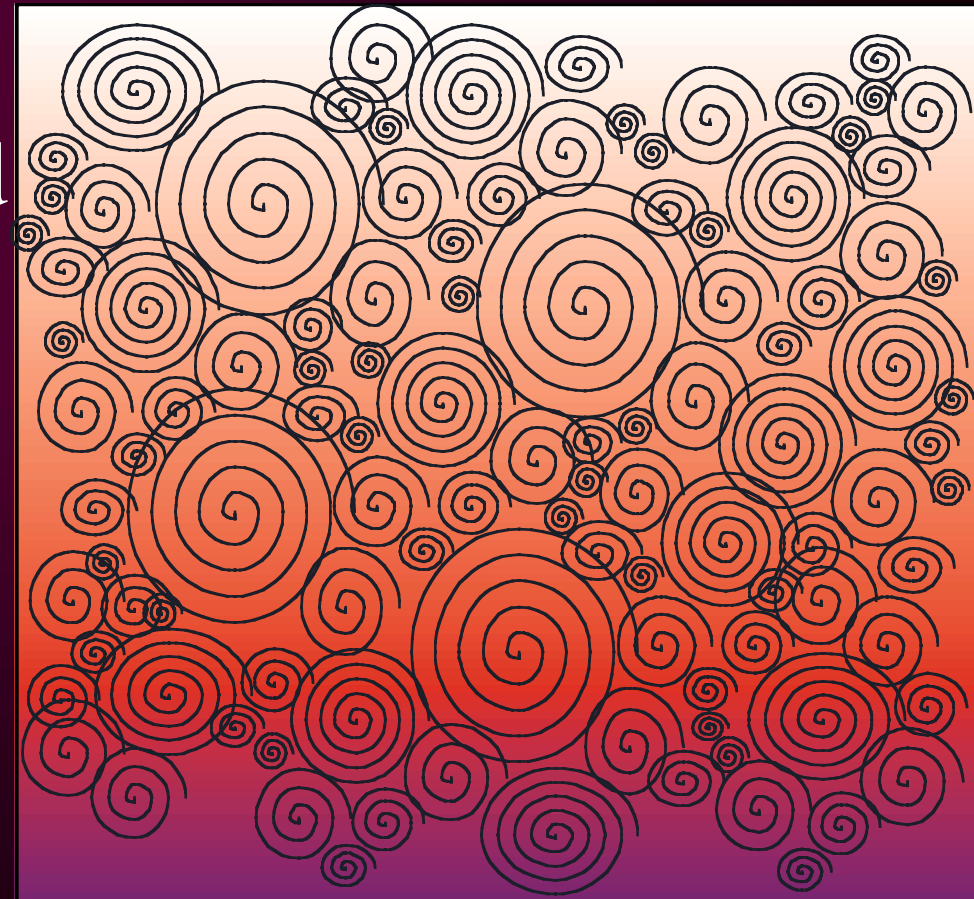
turbulence grows to flatten steep slopes in plasmas

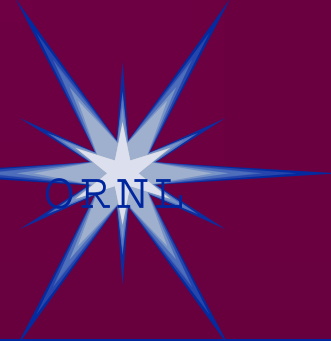
We need to control the turbulence

Turbulence moves things down the slope



The turbulent swirls
(eddies) move the heat and
density toward the edge



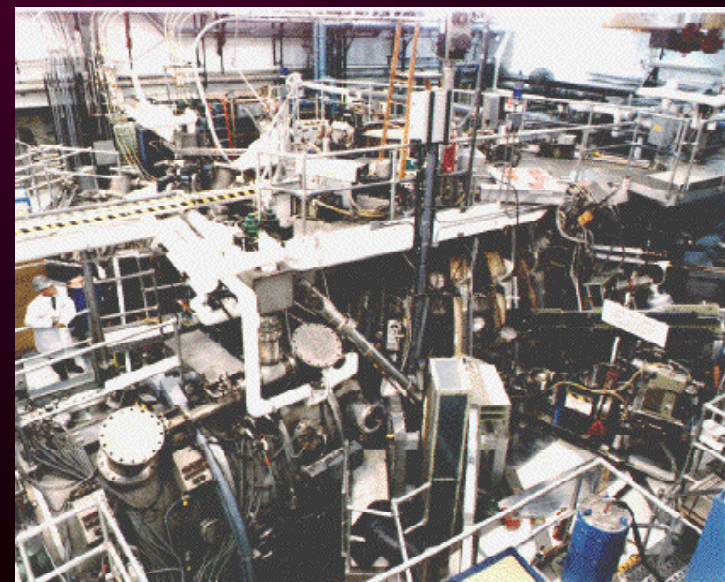
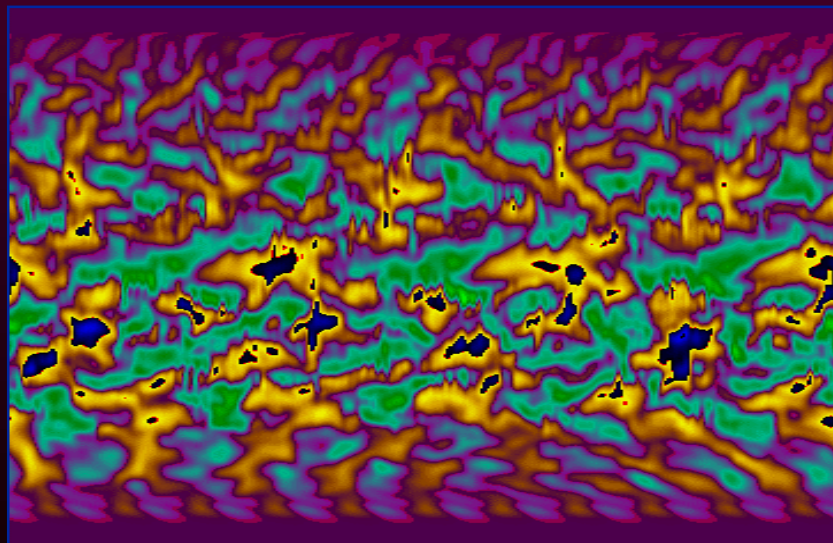
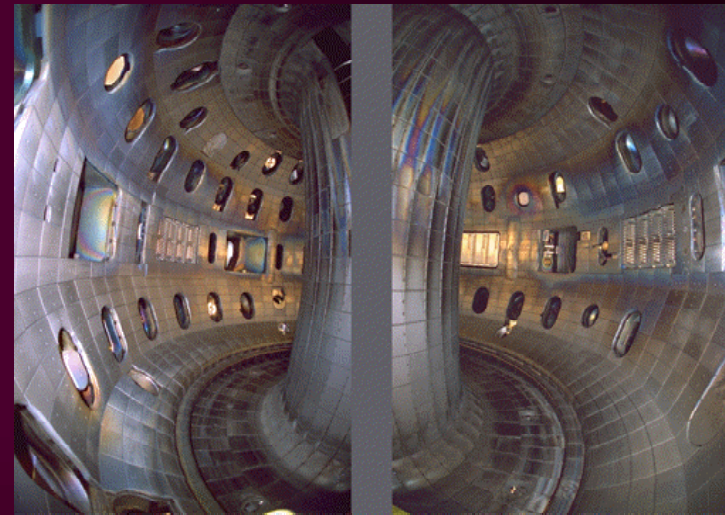


Challenges on the path to Fusion



Heating
Fueling
Confinement

Plasma physics is on the leading edge of technology





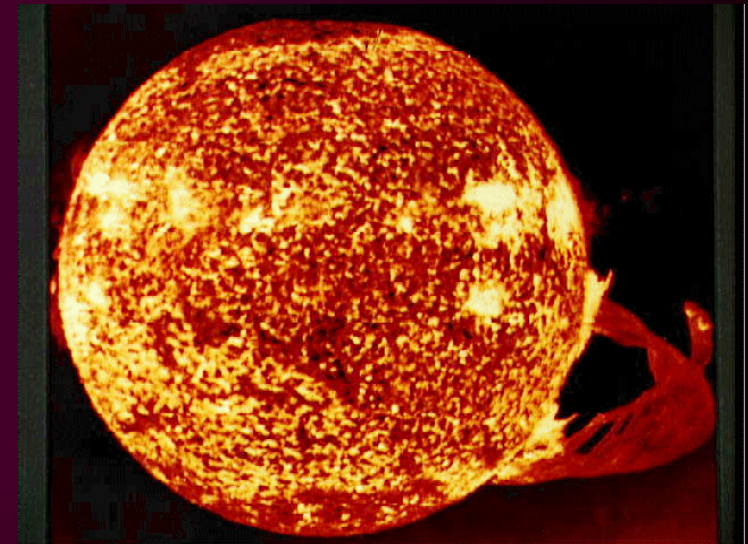
Turbulence is everywhere in nature



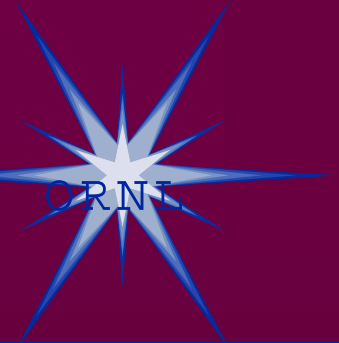
Turbulent transport is one of the main methods for relaxing gradients



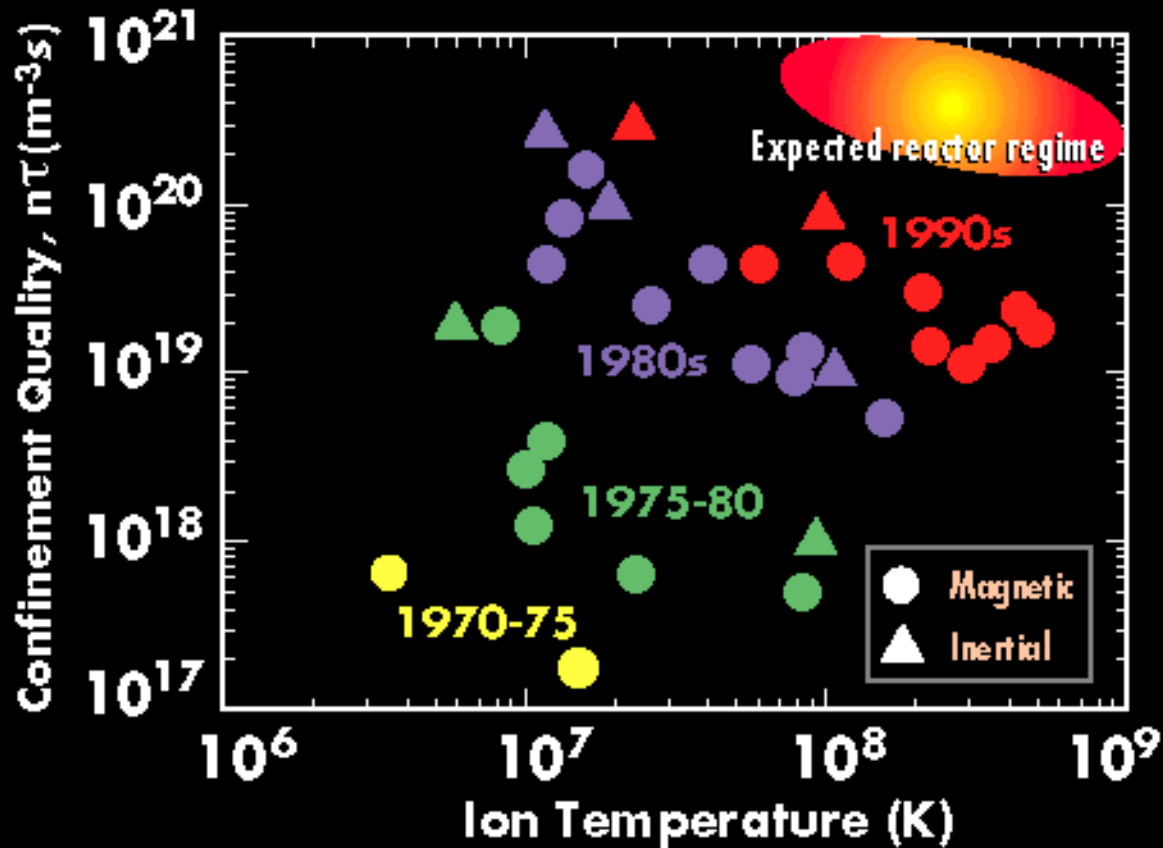
<ftp://mojave.wr.usgs.gov/pub/spurr/Spurr.html>



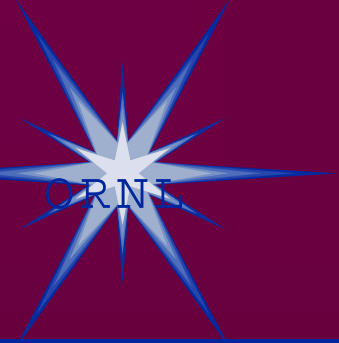
<http://info.pitt.edu/~maarten/work/soapjpps/dense.turb.JPG>



Progress towards fusion energy



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Summary



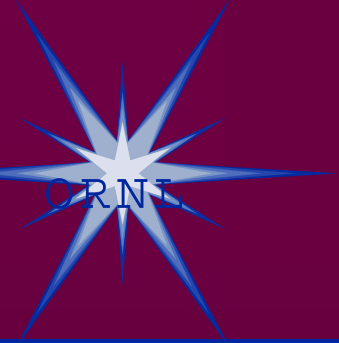
Plasmas are ubiquitous in our world

Science can indeed be both **fun** and **important**

Important discoveries and developments come from unexpected places so....

We must encourage people (especially young people) to explore the world around them

Stimulating interest stimulates learning which stimulates discovery and innovation which stimulates the economy



Web References



Fusion energy and plasma educational sites

http://FusionEd.gat.com/	General Atomics
http://FusEdWeb.pppl.gov/	Princeton Plasma Physics Laboratory
http://lasers.llnl.gov/lasers/education/ed.html	Lawrence Livermore National Laboratory
http://www.jet.uk/	Joint European Torus
http://www.ornl.gov/fed/fedhome.html/	Oak Ridge National Lab
http://www.ornl.gov/fed/theory/Theory_Home_page.html	
http://ffden-2.phys.uaf.edu	My home page at the Univ. of Alaska - Fairbanks

Astrophysics sites

http://umbra.nascom.nasa.gov/spd/	NASA Space Science
http://www.seds.org/billa/tnp/	The Nine Planets
http://www.stsci.edu:80/	Space Telescope Science Institute
http://bang.lanl.gov/solarsys/	Views of the Solar System
http://www.gi.alaska.edu/	Geophysical Institute (Aurora and Sprite info)
http://www.sec.noaa.gov/	NOAA Space weather site

Email me at: ffden@uaf.edu URL <http://ffden-2.phys.uaf.edu>